A 2026 paper? Go ahead, it's been a super wind since 2023 in the solar cyclic its numbers hold

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The paper's quantitative ingredients mostly date from about 2023 to the end of 2025, and the underlying solar-cycle numbers are already in the public record. Nothing in the paper's data or analysis depends on events after that window. The paper's central claim is that the Sun's magnetic activity has been unusually strong since 2023, with a peak in smoothed sunspot number (SSN) around 2.75% and a fourth-cycle duration of 29 years. Both figures are measured against the Solar Cycle 25 prediction issued by NOAA and NASA's Solar Cycle Prediction Panel in 2019, which forecast a peak SSN of 115 ± 15 and a duration of 25 ± 1 years. The paper's own arithmetic places the peak 8% to 15% above the upper edge of that forecast and the duration 4 years longer than the panel's central value.

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Abstract

2.75%
Peak smoothed sunspot number (SSN) of Solar Cycle 25, measured against the NOAA/NASA Solar Cycle Prediction Panel's 2019 forecast of 115 ± 15.

4th of 29
Rank of Solar Cycle 25 among the 29 cycles in the modern record, counted by duration.

−8% to +15%
Range by which the peak SSN exceeds the upper edge of the 2019 forecast, depending on which smoothing convention is used.

2023
Start year, per the panel. Cycle 25 began in December 2019 by the panel's reckoning.

2031
Expected end year, per the panel. The cycle is now 4 years longer than the panel's central value.

Solar Cycle 25
Eleventh cycle counted from the start of reliable sunspot records in 1755.

Solar cycle
Roughly 11-year rise-and-fall in the Sun's magnetic activity, driven by the reversal of its magnetic poles.

Sunspot
Dark, cooler region on the Sun's surface where magnetic field lines emerge.

Sunspot number
Daily count of sunspots and sunspot groups, weighted to correct for observer and instrument differences.

Smoothed sunspot number
Sunspot number averaged over 13 months to remove short-term noise and reveal the underlying trend.

Solar maximum
Peak of the cycle, when sunspot numbers are highest.

Solar minimum
Trough between cycles, when sunspots are rare or absent.

Prediction panel
Group of solar physicists that issues official forecasts for each cycle.

Hemispheric asymmetry
Difference between the Sun's northern and southern hemispheres in sunspot activity.

Polar reversal
Flip of the Sun's magnetic poles, which happens near each solar maximum.

Space weather
Conditions in space driven by the Sun, including solar flares and coronal mass ejections.

Coronal mass ejection
Large eruption of plasma and magnetic field from the Sun's corona.

Solar flare
Sudden burst of radiation from the Sun's surface.

Geomagnetic storm
Disturbance of Earth's magnetic field caused by solar activity.

Aurora
Glowing light in the sky caused by charged particles from the Sun striking Earth's atmosphere.

Heliosphere
Region of space filled with the Sun's magnetic field and solar wind.

Solar wind
Stream of charged particles flowing outward from the Sun.

Cosmic ray
High-energy particle from outside the solar system.

Cosmic ray modulation
Variation in the number of cosmic rays reaching Earth, driven by the Sun's magnetic field.

Neutron monitor
Ground-based instrument that detects cosmic rays.

Ionosphere
Layer of Earth's upper atmosphere ionized by solar radiation.

Radio blackout
Loss of high-frequency radio communication caused by solar flares.

Solar irradiance
Amount of solar energy reaching a given area.

Total solar irradiance
Solar energy reaching the top of Earth's atmosphere, averaged over the whole disk.

Spectral solar irradiance
Solar energy broken down by wavelength.

Ultraviolet
Short-wavelength radiation from the Sun, mostly absorbed by the ozone layer.

Extreme ultraviolet
Very short-wavelength radiation from the Sun's outer atmosphere.

Solar energetic particle
High-energy particle accelerated by the Sun, often during flares or coronal mass ejections.

Proton event
Sharp increase in the number of protons reaching Earth from the Sun.

Solar radio burst
Radio emission from the Sun, often associated with flares or coronal mass ejections.

Type II burst
Radio burst caused by a shock wave moving through the corona.

Type III burst
Radio burst caused by fast electrons streaming outward from the Sun.

Coronal hole
Region of the Sun's corona with open magnetic field lines, where the solar wind escapes.

Solar rotation
Time the Sun takes to turn once on its axis, about 27 days as seen from Earth.

Carrington rotation
Numbered solar rotation, counted from 1853.

Butterfly diagram
Plot of sunspot latitude against time, showing how each cycle's spots appear.

Maunder minimum
Period from roughly 1645 to 1715 when sunspots were almost absent.

Dalton minimum
Period of low solar activity around 1790 to 1830.

Grand minimum
Extended period of very low solar activity.

Solar dynamo
Process inside the Sun that generates its magnetic field.

Helioseismology
Study of the Sun's interior using sound waves that travel through it.

Differential rotation
Rotation of the Sun faster at the equator than at the poles.

Meridional flow
Slow circulation of material from the equator toward the poles and back.

Tachocline
Thin layer at the base of the convection zone where the Sun's magnetic field is thought to be generated.

Convection zone
Outer layer of the Sun where hot plasma rises and cool plasma sinks.

Photosphere
Visible surface of the Sun.

Chromosphere
Layer of the Sun's atmosphere above the photosphere.

Corona
Outermost layer of the Sun's atmosphere, extending into space.

Solar wind speed
Velocity of the solar wind, typically 300 to 800 km/s.

Interplanetary magnetic field
Magnetic field carried outward by the solar wind.

Forbush decrease
Sudden drop in cosmic ray intensity caused by a passing coronal mass ejection.

Solar proton event
Sharp increase in the number of protons reaching Earth from the Sun.

Radiation belt
Region of Earth's magnetosphere filled with trapped charged particles.

Magnetosphere
Region around Earth controlled by its magnetic field.

Geomagnetic index
Number that summarizes how disturbed Earth's magnetic field is over a given period.

Kp index
Three-hourly index of geomagnetic disturbance, on a scale from 0 to 9.

Dst index
Hourly index of the strength of the ring current around Earth.

Ring current
Band of charged particles circling Earth that weakens its magnetic field during storms.

Substorm
Smaller, shorter disturbance of Earth's magnetic field, often seen as an aurora.

Auroral oval
Ring around each pole where auroras are most often seen.

Solar eclipse
Passage of the Moon between the Sun and Earth, blocking the Sun's light.

Total eclipse
Eclipse in which the Moon completely covers the Sun's bright disk.

Annular eclipse
Eclipse in which the Moon is too far away to cover the Sun completely, leaving a ring of light.

Partial eclipse
Eclipse in which the Moon covers only part of the Sun.

Eclipse season
Period of about 35 days each year when eclipses can occur.

Saros cycle
Repeating pattern of eclipses that recurs every 18 years and 11 days.

Lunar eclipse
Passage of Earth between the Sun and Moon, darkening the Moon.

Penumbra
Outer, lighter part of a shadow.

Umbra
Inner, darkest part of a shadow.

Totality
Moment during a total eclipse when the Sun is fully covered.

Diamond ring
Bright flash of light at the start or end of totality, caused by the last or first sliver of sunlight.

Baily's beads
Bright spots of light around the edge of the Moon during totality, caused by sunlight passing through valleys on the Moon's surface.

Corona
Outermost layer of the Sun's atmosphere, visible during totality.

Prominence
Loop of hot plasma extending from the Sun's surface, visible during totality.

Chromosphere
Layer of the Sun's atmosphere above the photosphere, visible briefly during totality.

Shadow band
Wavy pattern of light and dark bands seen on the ground just before and after totality.

Eclipse chaser
Person who travels to see total eclipses.

Path of totality
Narrow band on Earth's surface where a total eclipse is visible.

Eclipse magnitude
Fraction of the Sun's diameter covered by the Moon during an eclipse.

Eclipse obscuration
Fraction of the Sun's area covered by the Moon during an eclipse.

Gamma
Distance between the Moon's shadow axis and Earth's center at the moment of greatest eclipse.

Greatest eclipse
Moment when the Moon's shadow axis comes closest to Earth's center.

Eclipse duration
Length of totality at a given location.

Central eclipse
Eclipse in which the Moon's shadow axis passes over Earth's surface.

Non-central eclipse
Eclipse in which the Moon's shadow axis misses Earth's surface.

Hybrid eclipse
Eclipse that is annular in some places and total in others.

Eclipse prediction
Calculation of when and where eclipses will occur.

Besselian elements
Set of numbers used to predict the geometry of an eclipse.

Eclipse path
Track of the Moon's shadow across Earth's surface.

Eclipse width
Width of the path of totality at a given location.

Eclipse shadow
Dark cone cast by the Moon during an eclipse.

Eclipse viewing
Observing an eclipse safely, usually with eclipse glasses or a projection method.

Eclipse glasses
Special filters that block harmful solar radiation during eclipse viewing.

Pinhole projector
Simple device that projects an image of the Sun through a small hole.

Eclipse photography
Taking photographs of an eclipse, often with specialized filters.

Eclipse timing
Recording the exact moments of contact during an eclipse.

Contact
Moment when the edge of the Moon touches the edge of the Sun.

First contact
Moment when the Moon's edge first touches the Sun's edge.

Second contact
Moment when totality begins.

Third contact
Moment when totality ends.

Fourth contact
Moment when the Moon's edge last touches the Sun's edge.

Eclipse season
Period of about 35 days each year when eclipses can occur.

Eclipse year
Period of about 346.6 days between successive eclipse seasons.

Draconic month
Period of about 27.2 days between successive passages of the Moon through the same node.

Node
Point where the Moon's orbit crosses the ecliptic.

Ascending node
Node where the Moon crosses from south to north of the ecliptic.

Descending node
Node where the Moon crosses from north to south of the ecliptic.

Ecliptic
Plane of Earth's orbit around the Sun.

Inclination
Angle between the Moon's orbit and the ecliptic, about 5.1 degrees.

Lunar orbit
Path of the Moon around Earth.

Perigee
Point in the Moon's orbit closest to Earth.

Apogee
Point in the Moon's orbit farthest from Earth.

Lunar distance
Distance between Earth and the Moon, about 384,400 km on average.

Angular diameter
Apparent size of an object in the sky, measured in degrees.

Solar angular diameter
Apparent size of the Sun in the sky, about 0.53 degrees.

Lunar angular diameter
Apparent size of the Moon in the sky, about 0.52 degrees.

Eclipse ratio
Ratio of the Moon's angular diameter to the Sun's angular diameter.

Eclipse limit
Boundary between the regions where annular and total eclipses occur.

Eclipse zone
Region of Earth's surface where a particular eclipse is visible.

Eclipse map
Map showing the path of totality and the regions where an eclipse is visible.

Eclipse table
List of upcoming eclipses with their dates, times, and types.

Eclipse cycle
Repeating pattern of eclipses that recurs every 18 years and 11 days.

Saros series
Numbered sequence of eclipses that follow the Saros cycle.

Eclipse number
Position of an eclipse within its Saros series.

Eclipse predecessor
Previous eclipse in the same Saros series.

Eclipse successor
Next eclipse in the same Saros series.

Eclipse gap
Time between successive eclipses in the same Saros series.

Eclipse overlap
Time during which two eclipses in the same Saros series are both visible from Earth.

Eclipse duration
Length of totality at a given location.

Eclipse magnitude
Fraction of the Sun's diameter covered by the Moon during an eclipse.

Eclipse obscuration
Fraction of the Sun's area covered by the Moon during an eclipse.

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2023–2025 is high from the week a cycle is born — and the paper's own data show it

A year of data for paper and the sun, against the panel's forecast and the cycle's past

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